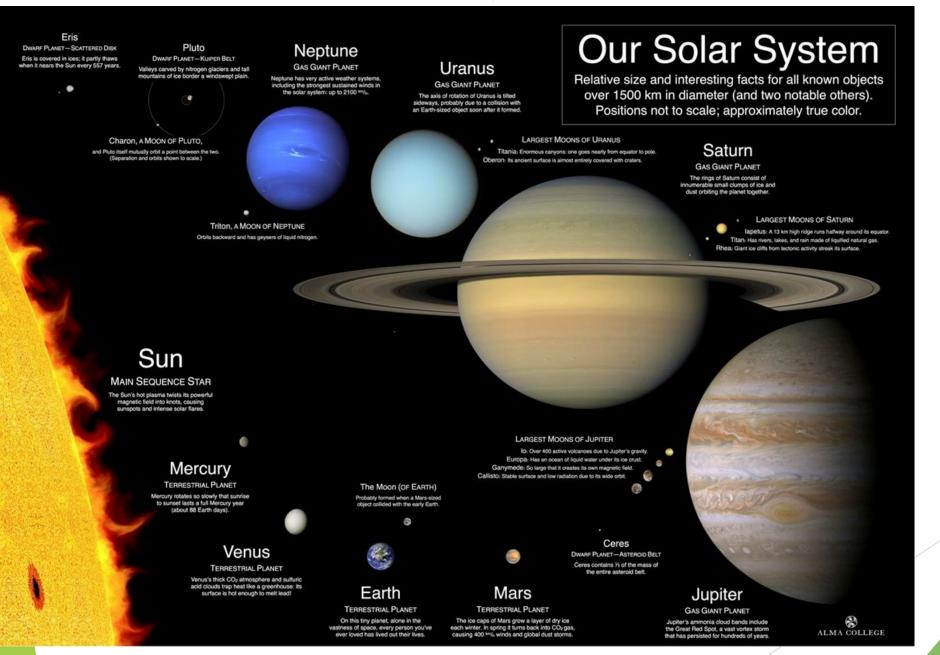
To explore the solar system – A case study of Callisto

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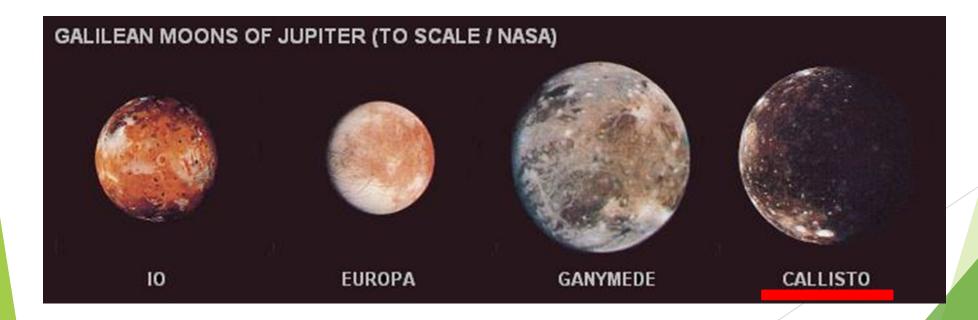
Introduction: solar system



• The solar system formed about 4.6 billion years ago.

Introduction: Callisto

- Callisto is the fourth Galilean moon orbits around Jupiter.
- Callisto formed at the same time as the solar system.
- The size of Callisto compared with other Galilean moons:



Introduction: Callisto

- Mass: 1.07×10²³ kg (0.018 earth mass)
- Mean radius: 2410.3 km (0.378 Earth radius)
- Mean density: 1.83 g/cm³ (mixture of ices and rocks)



Introduction: Callisto

• Exosphere:

Carbon dioxide is revealed by Galileo spacecraft, and molecular oxygen is revealed by Hubble space telescope.

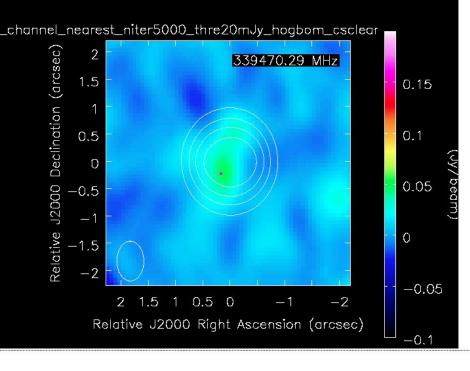
- Solid surface: water ice, carbon dioxide, silicates, and organic compounds
- Interior structure:

Callisto may have subsurface oceans of liquid water.

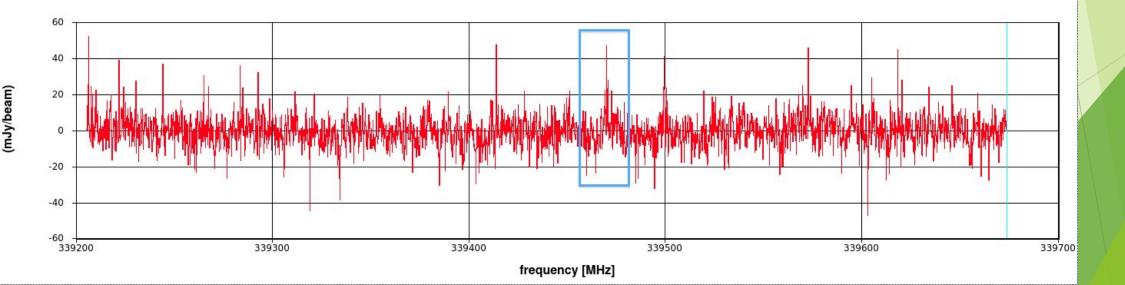
Observations

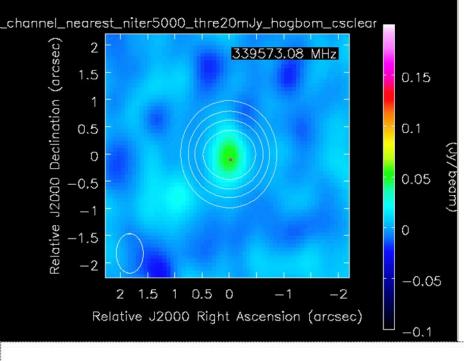
- The observational data are taken from ALMA (Atacama Large Millimeter/submillimeter Array) that is the largest ground-based astronomical facilities in the world.
- ALMA is an interferometer array in Chile and it used 23 antennas to observe Callisto in our data (band 7).
- Callisto was observed on 23-Oct-2012.
- Total integration time: 206.64 seconds



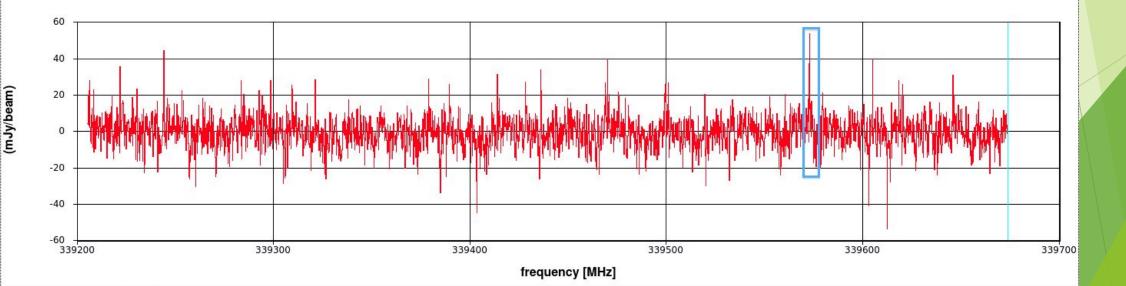


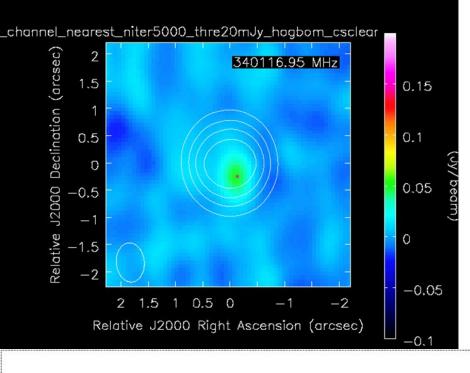
- spectral line s1
- sky frequency: 339470.290 MHz
- rms: 8.8mJy/beam
- S/N: 5.4 σ





- spectral line s2
- sky frequency: 339573.080 MHz
- rms: 8.8mJy/beam
- S/N: 6.1 σ

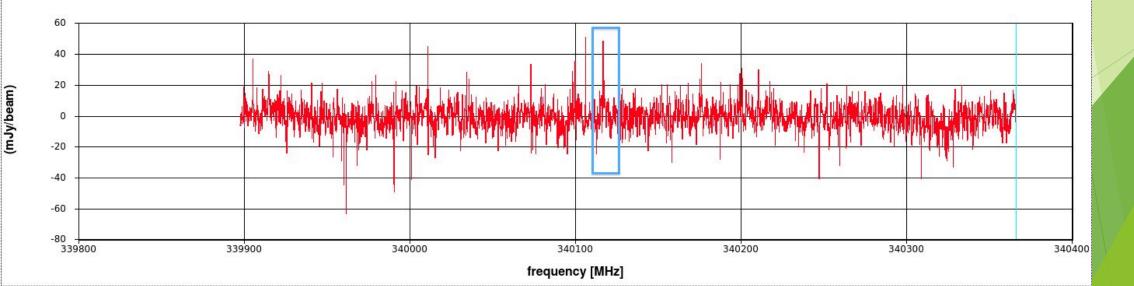


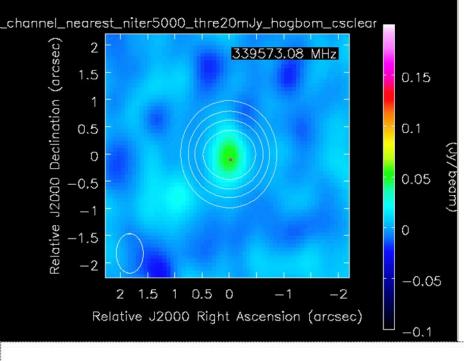


spectral line s3

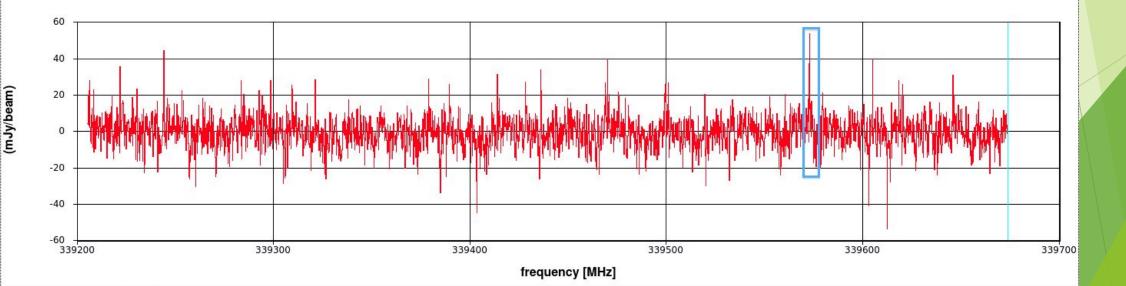
- sky frequency: 340116.950 MHz
- rms: 8.9 mJy/beam

• S/N: 5.6 σ



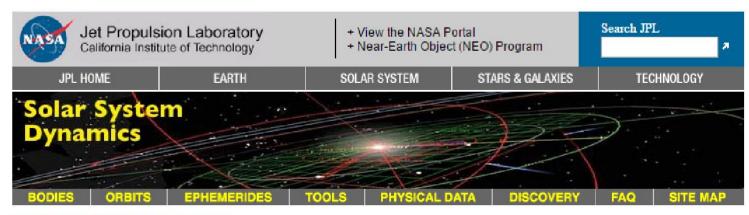


- spectral line s2
- sky frequency: 339573.080 MHz
- rms: 8.8mJy/beam
- S/N: 6.1 σ



Observations

• Doppler shift need to be considered.



HORIZONS Web-Interface

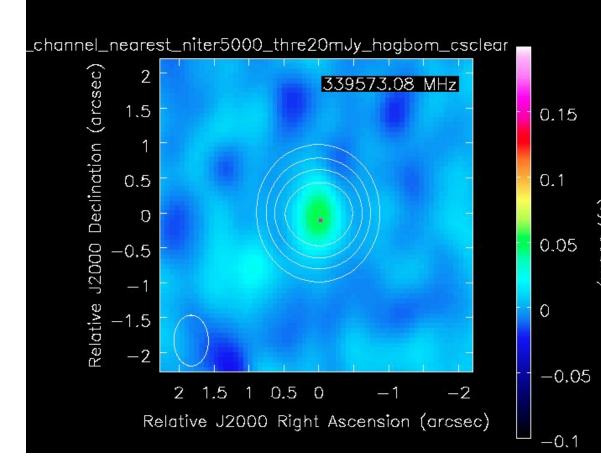
This tool provides a web-based *limited* interface to JPL's HORIZONS system which can be used to generate ephemerides for solar-system bodies. Full access to HORIZONS features is available via the primary telnet interface. HORIZONS system news shows recent changes and improvements. A web-interface tutorial is available to assist new users.

Current Settings

Ephemeris Type [change] : OBSERVER Target Body [change] : Callisto (JIV) [504] Observer Location [change] : Atacama Large mm/sub-mm Array (ALMA) Center of Array [-7] (292°14'43.1"E, 23°01'45.0"S, 5079.71 m) Time Span [change] : Start=2012-10-23, Stop=2012-10-24, Step=1 m Table Settings [change] : defaults Display/Output [change] : default (formatted HTML)

*****	*****	*****	****	***	*****	*****	***	*****	******	******	***	******	**********	*******	****	*****	ł
Date	_(UT)_	HR:MN	R.A	۱((ICRF/	J 2000	.0)	DEC	APmag	S-brt		delta	deldot	S-0-T	/r	S-T-(5
*****	*****	********	****	: sk: sk: 3	*****	*****	ak ak a	*****	*****	******	***:	*****	*****	******	****	*****	k
2012-	Oct-23	07:12	04	58	26.11	+21	50	29.1	5.95	6.63	4.3	28666981991501	-24.9207826	134.3610	/L	8.1289	5
2012-	Oct-23	07:13	04	58	26.09	+21	50	29.1	5.95	6.63	4.3	28665982534677	-24.9174479	134.3618	/L	8.1284	4
2012-	Oct-23	07:14	04	58	26.07	+21	50	29.0	5.95	6.63	4.3	28664983211620	-24.9141125	134.3626	/L	8.1283	3
2012-	Oct-23	07:15	04	58	26.05	+21	50	29.0	5.95	6.63	4.3	28663984022360	-24.9107763	134.3634	/L	8.1282	2

Observations



Sky frequency: 339573.080 MHz

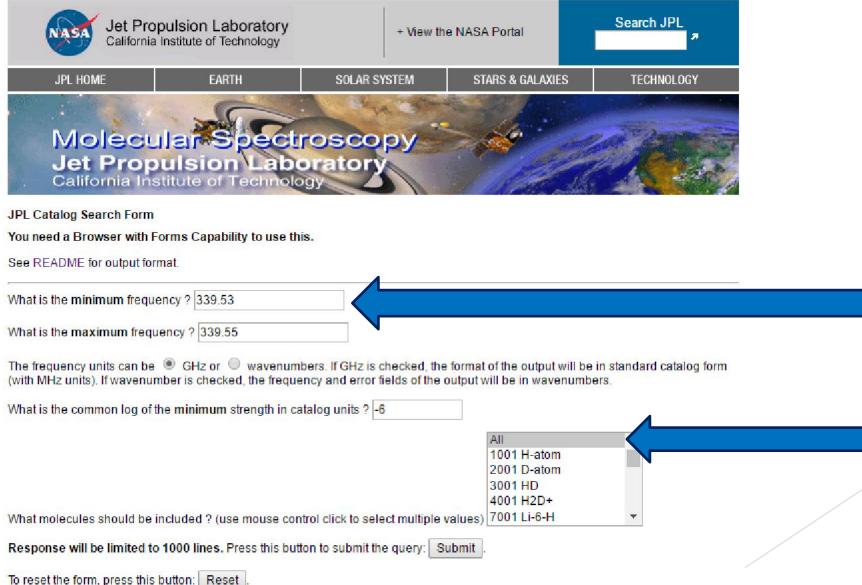
Doppler effect: $\Delta f = f_{sky} - f_{rest}$

f_{sky}: sky frequency f_{rest}: rest frequency v: velocity of the receiver relative to the source c: speed of light

 $\Delta f=(-v/c) f_{sky}$

Rest frequency: 339544.854MHz

To search candidates from Jet Propulsion Laboratory.



37002 C3H

37002 C3H									
339540.3431839.2272 -4.2770 2 27.0947 11 370021325 5 0 1 6 5 6-1 0 6									
42003 NH2CN									
339544.9830 0.0088 -2.1501 3 341.6674105 42003140417 512 0 16 511 0									
339544.9830 0.0088 -2.1501 3 341.6674105 42003140417 513 0 16 512 0									
46004 C2H5OH									
339544.0950 0.0500 -4.3164 3 52.3989 17 -460041404 8 7 1 2 8 6 2 2									
339544.0950 0.0500 -4.3164 3 52.3989 17 -460041404 8 7 2 2 8 6 3 2									
46011 DOCO+									
339534.7258 98.3990 -4.8934 3 1403.6907 35 46011 3031710 7 1610 6									
339534.7258 98.3990 -4.8934 3 1403.6907 35 46011 3031710 8 1610 7									
52009 CH3C1-37									
339530.3239 0.0161 -4.5914 3 697.8120 56 52009 213131114 121113									
339531.0431 0.0166 -4.6252 3 697.8119 52 52009 213131113 121112									
339535.2245 0.0183 -4.5577 3 697.8114 60 52009 213131115 121114									
339535.9348 0.0188 -4.6590 3 697.8113 48 52009 213131112 121111									
53001 C2H3CN									
339547.1295 0.9222 -5.3050 3 667.8561 91 53001140544 441 245 44 342 24									
339547.1322 0.9222 -5.3245 3 667.8561 87 53001140544 441 243 44 342 243									
339547.1632 0.9222 -5.3149 3 667.8560 89 53001140544 441 244 44 342 244									
55001 C2H5CN									
339544.3560 0.0500 -3.9303 3 404.4402 99 -55001 30349 842 49 743									
57003 HCCCH20D									
339530.0111999.9999 -5.3183 3 856.5739 77 570031404382810 0 3728 9 0									
339530.0111999.9999 -5.3183 3 856.5739 77 570031404382811 0 372810 0									
339537.3216999.9999 -5.4815 3 911.8266 77 5700314043829 9 1 3729 8 1									
339537.3216999.9999 -5.4815 3 911.8266 77 570031404382910 1 3729 9 1									
60003 CH3OCHO									
339548.4026 0.0090 -5.1865 3 484.5465150 600031404371325 3 371226 3									

rest frequency: 339544.854MHz

• Selection criteria:

1. Integrated intensity can't be too low.

2. Frequencies can't be more or less than rest frequency by2MHz(0.107km/s/channel).

• From database to spectral lines

Jet Propulsi California Institu	on Laboratory ite of Technology	+ View the NASA Portal	Search JPL	Frequence	uency range	e:
JPL HOME	EARTH SOL	AR SYSTEM STARS & GAL/	AXIES TECHNOLOGY	330 2	206 GHZ to	339.674 GHz
A States		A State of the		000.2		000.074 0112
Molecula	r Spectroso	copy 💦 💉	de ter	-T		
Jet Propuls California Institut	sion Laborato	orv /	GCA AN			
JPL Catalog Search Form				<i>6</i>		
You need a Browser with Forms	Capability to use this.					
See README for output format.						
What is the minimum frequency ?	339.2	/				
What is the maximum frequency	? 339.7					
		iz is checked, the format of the outp	ut will be in standard catalog for	~		
(with MHz units). If wavenumber is	s checked, the frequency and e	rror fields of the output will be in w	avenumbers.			
What is the common log of the mi	nimum strength in catalog unit	s ? <mark>-</mark> 6				
		41010 CF 41011 K4	I3CN v8=1			
		42001 CH	13CN-15			
		42002 CH 42003 NH	12CN			
What molecules should be include	ed ? (use mouse control click t	o select multiple values) 42004 Ca	D			
Response will be limited to 1000	lines. Press this button to sub	mit the query: Submit .				
				/		

42003 NH2CN

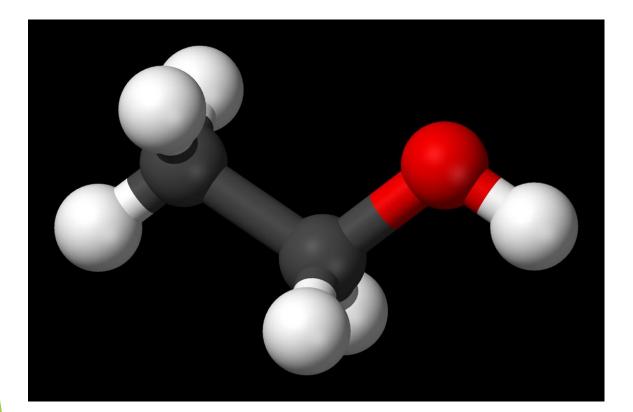
	42000 NHZCN				
339238.0406	0.0055 -1.7988 3	179.3581105	42003140417 215 1	16 214 1	no signal
339239.4169	0.0104 -2.7304 3	385.0814 35	42003140417 512 1	16 511 1	no signal
339239.4169	0.0104 -2.7304 3	385.0814 35	42003140417 513 1	16 512 1	
339347.0172	0.0078 -2.0553 3	297.0513105	42003140417 414 1	16 413 1	
339347.0186	0.0078 -2.0553 3	297.0513105	42003140417 413 1	16 412 1	
339376.4504	0.0111 -2.8750 3	451.5164 35	42003140417 611 0	16 610 0	
339376.4504	0.0111 -2.8750 3	451.5164 35	42003140417 612 0	16 611 0	no signal
339422.7962	0.0064 -2.3787 3	228.4772 35	42003140417 315 1	16 314 1	5
339423.4774	0.0064 -2.3787 3	228.4773 35	42003140417 314 1	16 313 1	
339450.5912	0.0063 -1.7026 3	140.1788105	42003140417 017 1	16 016 1	
339544.9830	0.0088 -2.1501 3	341.6674105	42003140417 512 0	16 511 0	
339544.9830	0.0088 -2.1501 3	341.6674105	42003140417 513 0	16 512 0	
339692.1611	0.0070 -2.4243 3	251.5365 35	42003140417 414 0	16 413 0	
339692.1632	0.0070 -2.4243 3	251.5365 35	42003140417 413 0	16 412 0	

Spectral line candidates

- 17 candidates for spw0 (NH2CN, C2H5OH, C2H5CN, etc.)
- 10 candidates for spw1 (CH3OCHO, C2H3CN, CH3CHO, etc.)
- 4 candidates for spw2 (HCOCH2OH, HNO3, Phenol, HSC)
- 11 candidates for spw3 (HO3, HOCI, C2H3CN, etc.)

Spectral line candidates







46004 C2H5OH 339201.6420 0.0500 -3.9171 3 84.5342 27 -46004140413 7 6 2 13 6 7 2 339201.6420 0.0500 -3.9171 3 84.5342 27 -46004140413 7 7 2 13 6 8 2	Line E1
339235.7497 0.0109 -3.6160 3 158.4519 41 46004140420 218 1 19 217 1	Line E2
339282.5580 0.0044 -4.1782 3 111.1897 31 46004140415 412 0 14 312 1	Line E3
339312.57000.0500 -3.9566 376.9355 25 -46004140412 7 5 212 6 6 2339312.57000.0500 -3.9566 376.9355 25 -46004140412 7 6 212 6 7 2	Line E4
339398.45100.0500 - 4.0081 369.9230 23 - 46004140411 7 4 211 6 5 2339398.45100.0500 - 4.0081 369.9230 23 - 46004140411 7 5 211 6 6 2	Line E5
339423.4340 0.0500 -3.5527 3 292.2965 65 -46004140432 131 2 32 032 2	Line E6
339453.9700 0.0500 -3.9583 3 360.6054 71 -46004140435 233 2 35 134 2	Line E7
339461.7880 0.0500 -3.6075 3 154.6536 41 -46004140420 219 1 19 218 1	Line E8
339463.30900.0500 - 4.0767 363.4963 21 - 46004140410 7 3 210 6 4 2339463.30900.0500 - 4.0767 363.4963 21 - 46004140410 7 4 210 6 5 2	Line E9
339510.73700.0500 -4.1720 357.6551 19 -460041404 9 7 2 29 6 3 2339510.73700.0500 -4.1720 357.6551 19 -460041404 9 7 3 29 6 4 2	Line E10
339544.09500.0500 -4.3164 352.3989 17 -460041404 8 7 1 28 6 2 2339544.09500.0500 -4.3164 352.3989 17 -460041404 8 7 2 28 6 3 2	Line E11
339566.32100.0500 -4.5829 347.7274 15 -460041404 7 7 0 27 6 1 2339566.32100.0500 -4.5829 347.7274 15 -460041404 7 7 1 27 6 2 2	Line E12
339577.8590 0.0500 -3.8542 3 292.2966 65 -46004140432 231 2 32 132 2	Line E13

Ethanol

Summary

- Callisto is an important astronomical object because there may be organic compounds on its surface and Callisto may be habitable.
- We need more data to prove that there is ethanol in exosphere of Callisto.

Implications

- Exosphere: chemical composition
- Cryovolcanos: a volcano that erupts ice-volcanic melt such as water, methane or ammonia.
- Subsurface oceans: molecules can provide evidence that Callisto have subsurface oceans.

The end